



Development of Laser Sources and Diagnostics for Probing Relativistic Laser-Matter Interaction

Master thesis

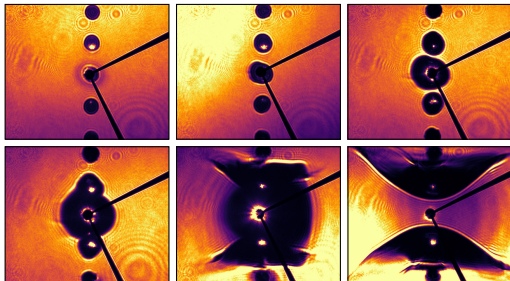
Martin Beyer

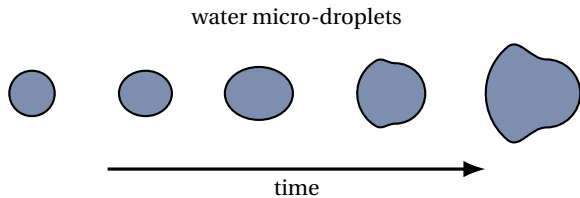
supervisors:

Prof. Dr. Malte Kaluza

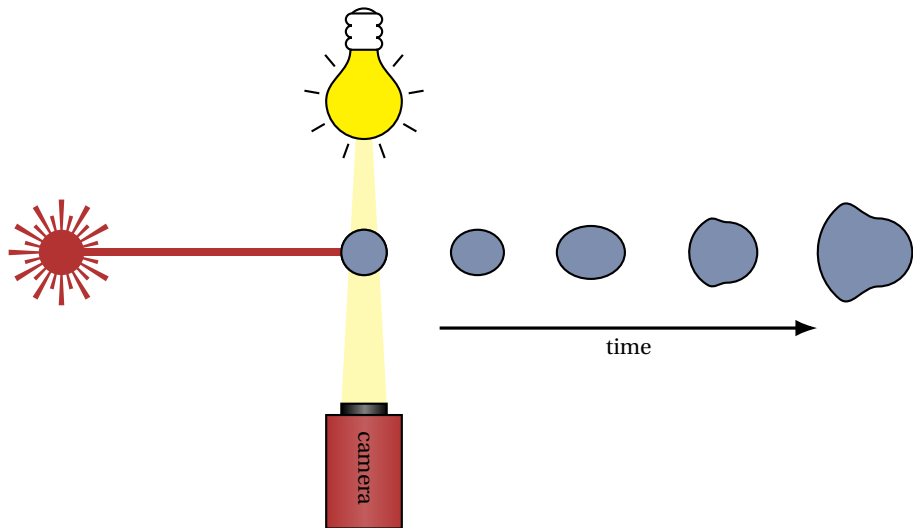
Dr. Yasmina Azamoum

December 9th, 2022

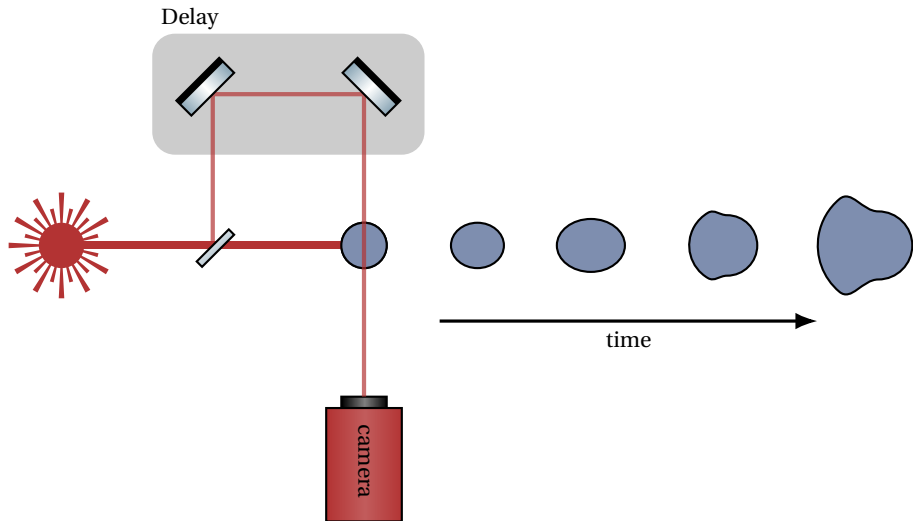




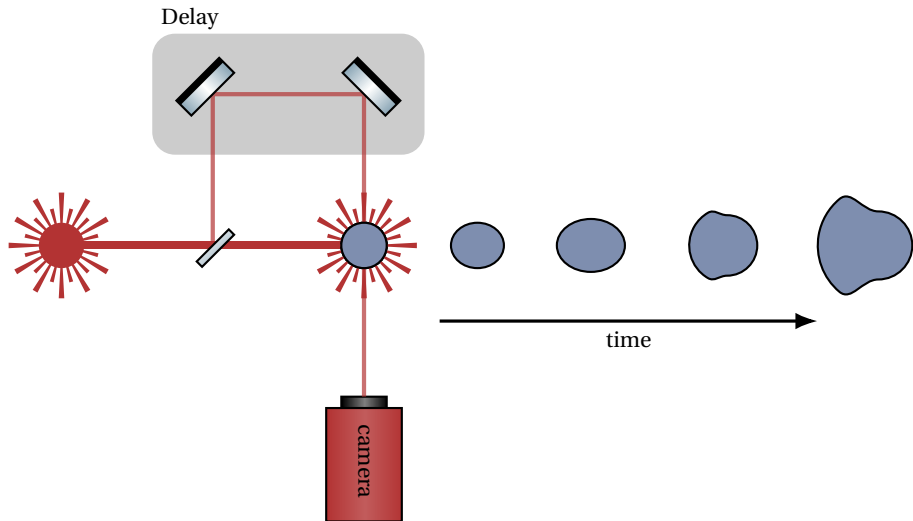
The Dilemma

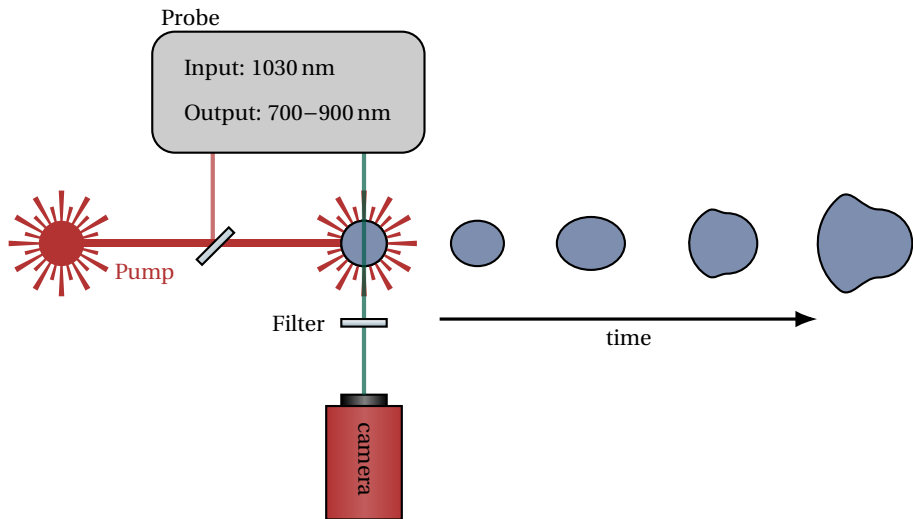


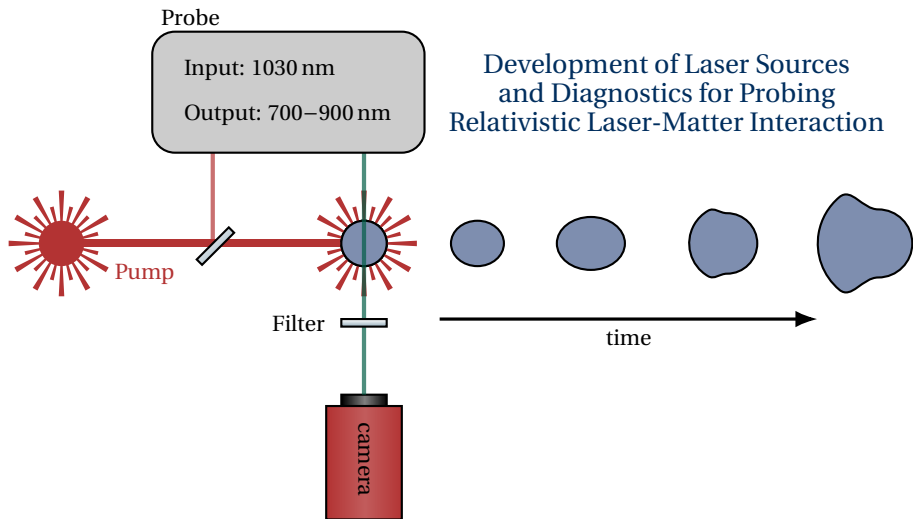
The Dilemma



The Dilemma







1. Fundamentals

- Laser-induced plasmas
- Optical parametric amplification

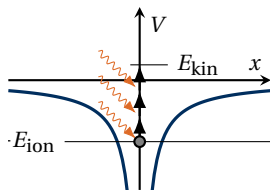
2. Probing relativistic laser-plasma interactions

3. SPM-based laser pulse generation

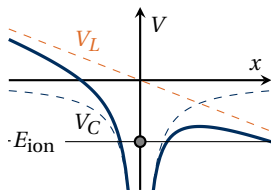
- Plasma: "A quasineutral gas of charged and neutral particles which exhibits collective behavior" [Chen, Introduction to plasma physics and controlled fusion]

- Plasma: "A quasineutral gas of charged and neutral particles which exhibits collective behavior" [Chen, Introduction to plasma physics and controlled fusion]
- classical ionization: photoelectric effect $E = \hbar\omega > E_{\text{ion}}$,
water: $E_{\text{ion}} = 10.1 \text{ eV}$ vs. POLARIS: $E_{\text{photon}} = 1.2 \text{ eV}$

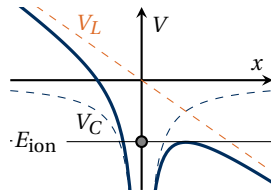
- Plasma: "A quasineutral gas of charged and neutral particles which exhibits collective behavior" [Chen, Introduction to plasma physics and controlled fusion]
- classical ionization: photoelectric effect $E = \hbar\omega > E_{\text{ion}}$,
water: $E_{\text{ion}} = 10.1 \text{ eV}$ vs. POLARIS: $E_{\text{photon}} = 1.2 \text{ eV}$
- For **higher** intensities:



(a) Multi-Photon Ionization



(b) Tunnel Ionization



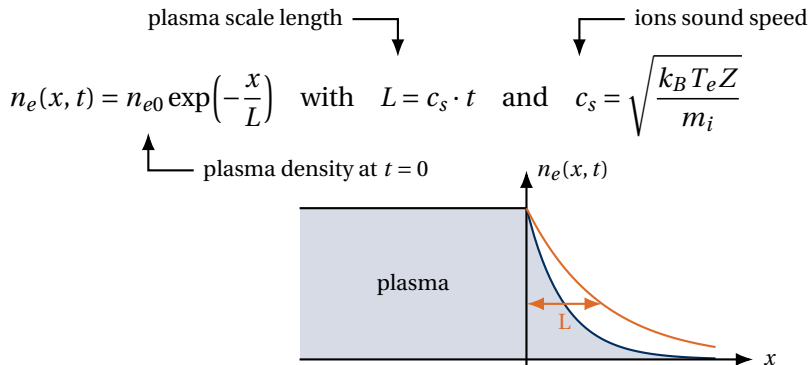
(c) Over the Barrier Ionization

- **plasma frequency:** $\omega_p^2 = \frac{n_e e^2}{\epsilon_0 m_e}$,

- plasma frequency: $\omega_p^2 = \frac{n_e e^2}{\epsilon_0 m_e}$, **critical density:** $n_{\text{cr}} = \frac{\omega^2 \epsilon_0 m_e}{e^2}$
⇒ plasma is opaque for wavelengths above $\lambda = 2\pi c / \omega_p$

- plasma frequency: $\omega_p^2 = \frac{n_e e^2}{\epsilon_0 m_e}$, critical density: $n_{cr} = \frac{\omega^2 \epsilon_0 m_e}{e^2}$
 \Rightarrow plasma is opaque for wavelengths above $\lambda = 2\pi c / \omega_p$

- Plasma expansion:**



(a)

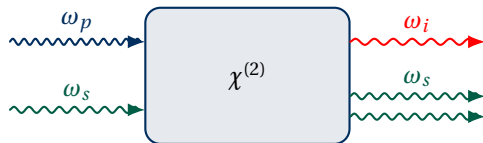


Fig.: Process of Difference frequency generation (DFG).

(a) Geometry. conversion of the pump photon ω_p to ω_s and $\omega_i = \omega_p - \omega_s$.

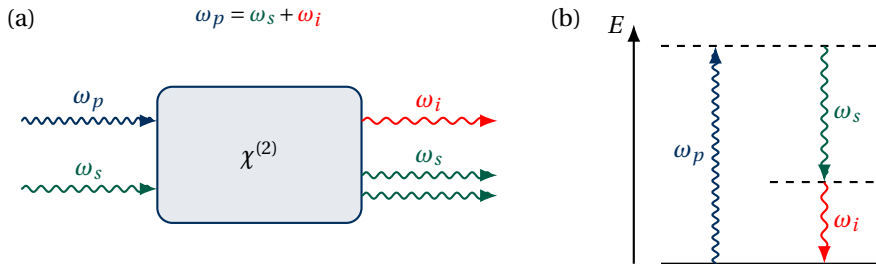


Fig.: Process of Difference frequency generation (DFG).

(a) Geometry. conversion of the pump photon ω_p to ω_s and $\omega_i = \omega_p - \omega_s$.

(b) Energy-level description.

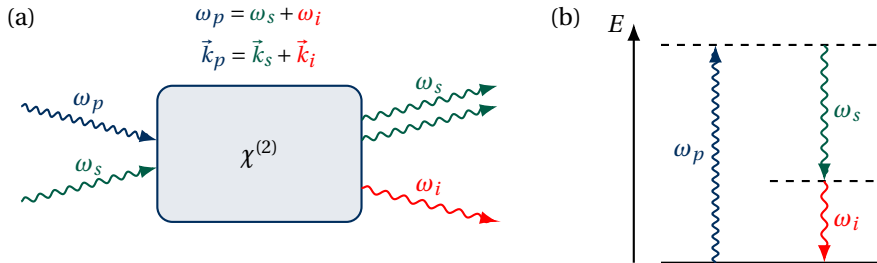


Fig.: Process of Difference frequency generation (DFG).

(a) Geometry. conversion of the pump photon ω_p to ω_s and $\omega_i = \omega_p - \omega_s$.

(b) Energy-level description.

1. Fundamentals

2. Probing relativistic laser-plasma interactions

- NOPA-based laser pulse generation
- Experimental setup
- Shadowgraphy of the plasma evolution

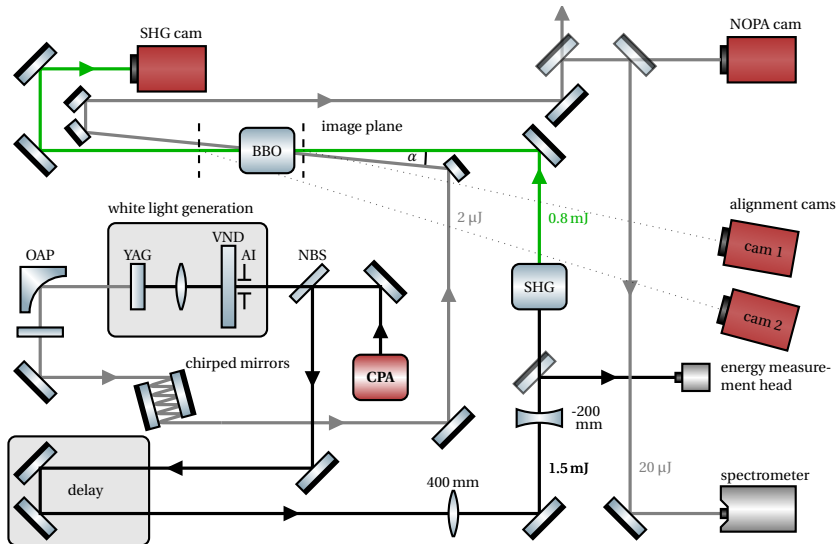
3. SPM-based laser pulse generation

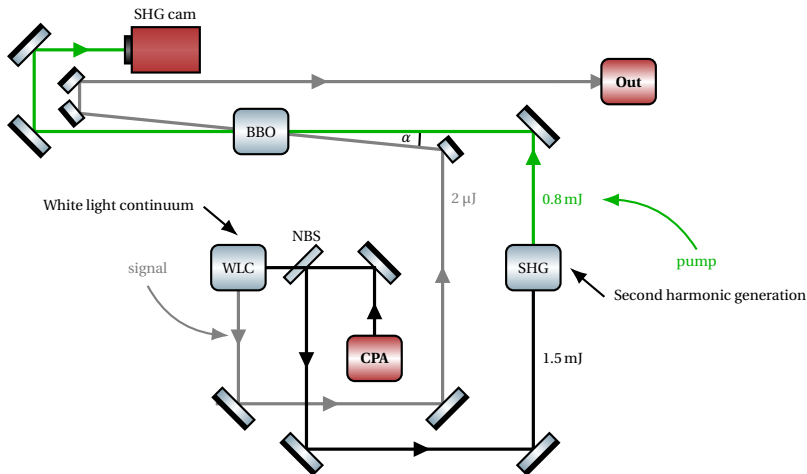
- **Pulse duration** $\tau < 130$ fs
⇒ high temporal resolution

- **Pulse duration** $\tau < 130$ fs
⇒ high temporal resolution
- **Wavelength** $\lambda = 800$ nm
⇒ No overlap with fundamental (1030 nm) and SH (515 nm) of the POLARIS laser

- **Pulse duration** $\tau < 130$ fs
⇒ high temporal resolution
- **Wavelength** $\lambda = 800$ nm
⇒ No overlap with fundamental (1030 nm) and SH (515 nm) of the POLARIS laser
- **Spectral bandwidth** $\Delta\lambda > 100$ nm
⇒ Support femtosecond pulse durations, enable spectrally resolved probing options

- **Pulse duration** $\tau < 130$ fs
⇒ high temporal resolution
- **Wavelength** $\lambda = 800$ nm
⇒ No overlap with fundamental (1030 nm) and SH (515 nm) of the POLARIS laser
- **Spectral bandwidth** $\Delta\lambda > 100$ nm
⇒ Support femtosecond pulse durations, enable spectrally resolved probing options
- **Energy** $E > 10\mu\text{J}$
⇒ proper illumination of the interaction region without target ionization





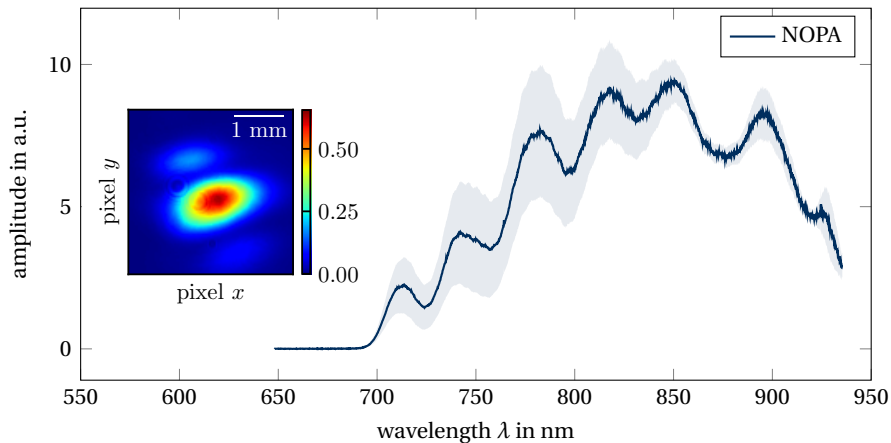
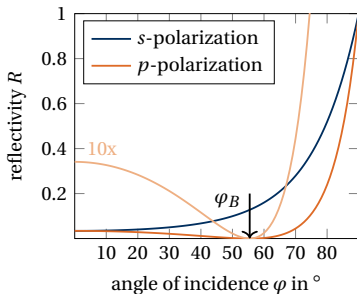


Fig.: Measured averaged spectrum (for 60 consecutive shots) of the NOPA. The standard deviation is indicated with the shaded area. Inset: Spatial profile of the NOPA.

- **Suppression of the plasma emission and scattered laser light**
 - ⇒ **spectral filter (bandpass centered at 800 nm)**

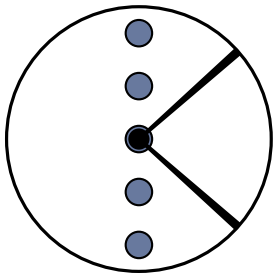
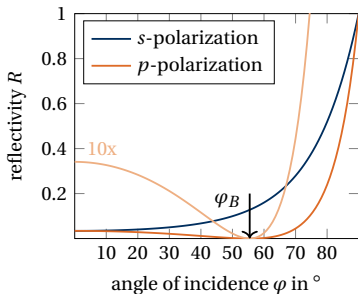
- **Suppression of the plasma emission and scattered laser light**

- ⇒ spectral filter (bandpass centered at 800 nm)
- ⇒ **polarization filter**

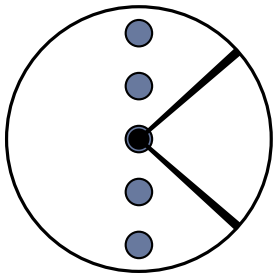
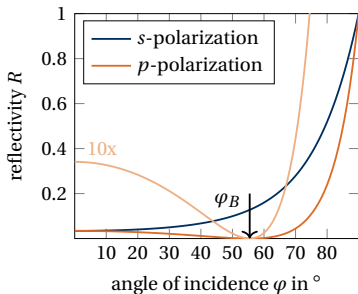


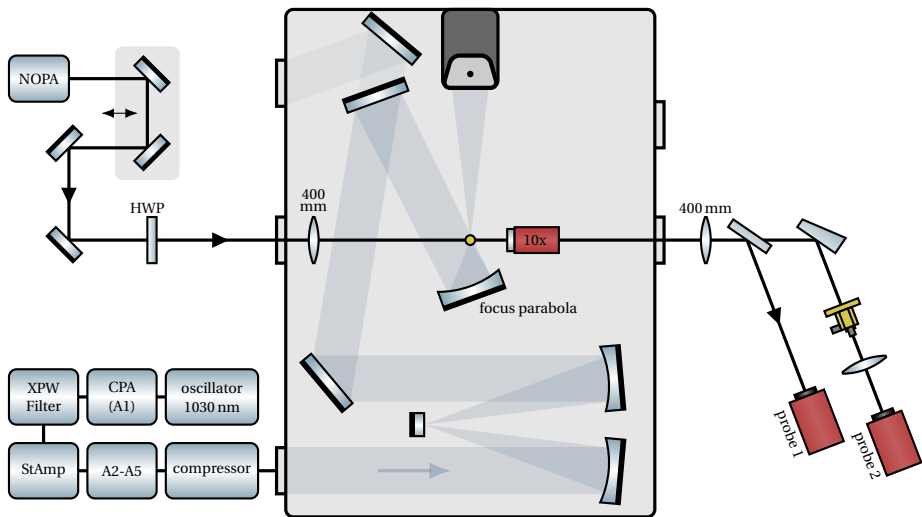
- **Suppression of the plasma emission and scattered laser light**

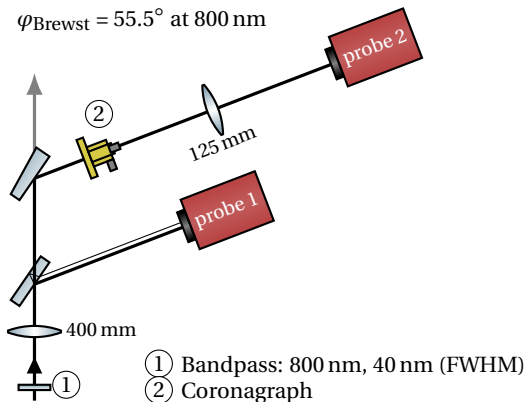
- ⇒ spectral filter (bandpass centered at 800 nm)
- ⇒ polarization filter
- ⇒ **spatial filter**

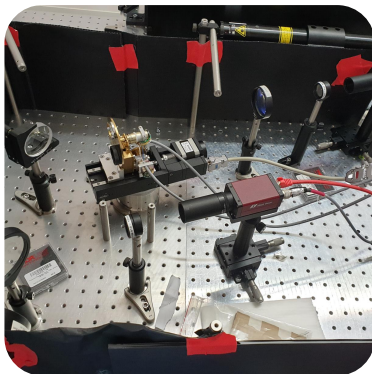
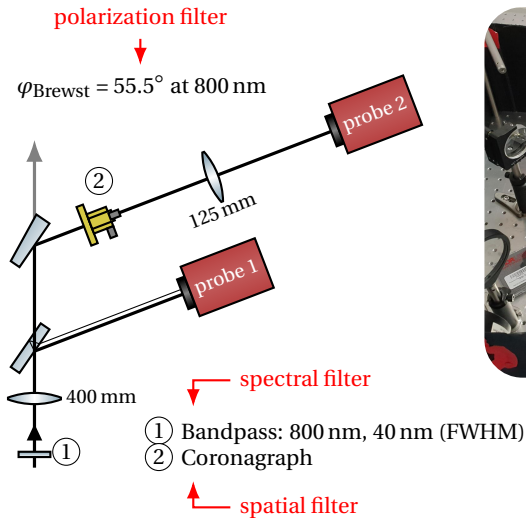


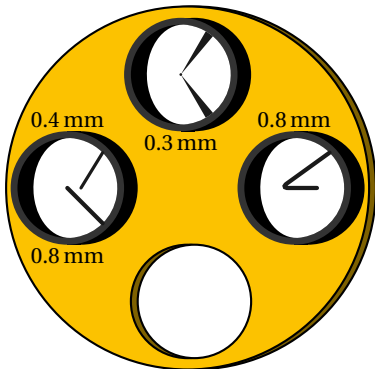
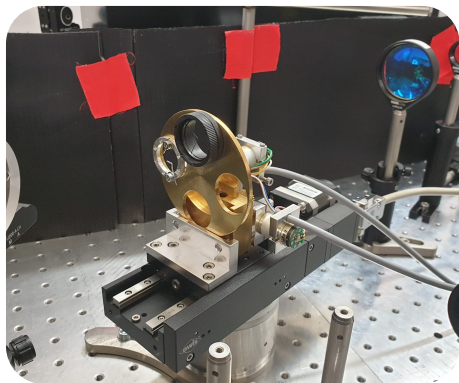
- **Suppression of the plasma emission and scattered laser light**
 - ⇒ spectral filter (bandpass centered at 800 nm)
 - ⇒ polarization filter
 - ⇒ spatial filter
- **resolution** ⇒ microscope objective

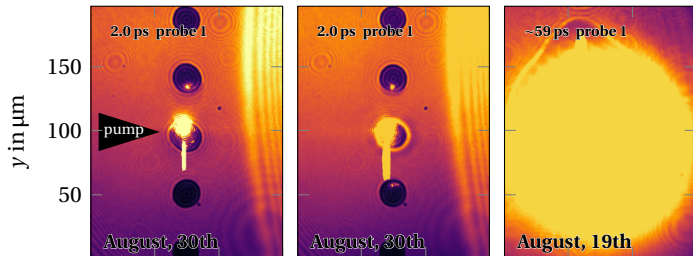


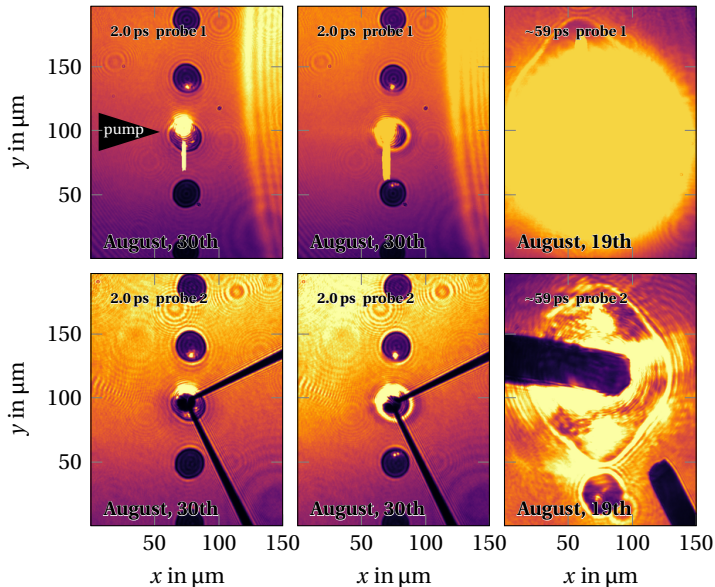




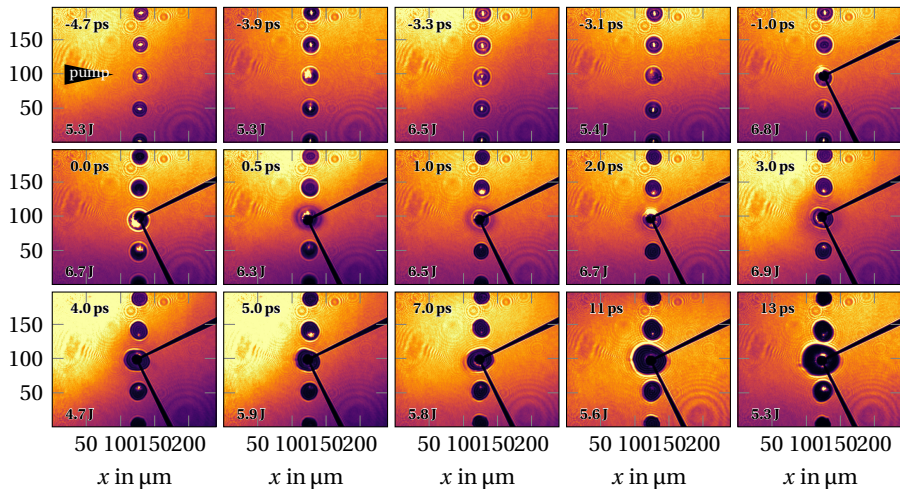




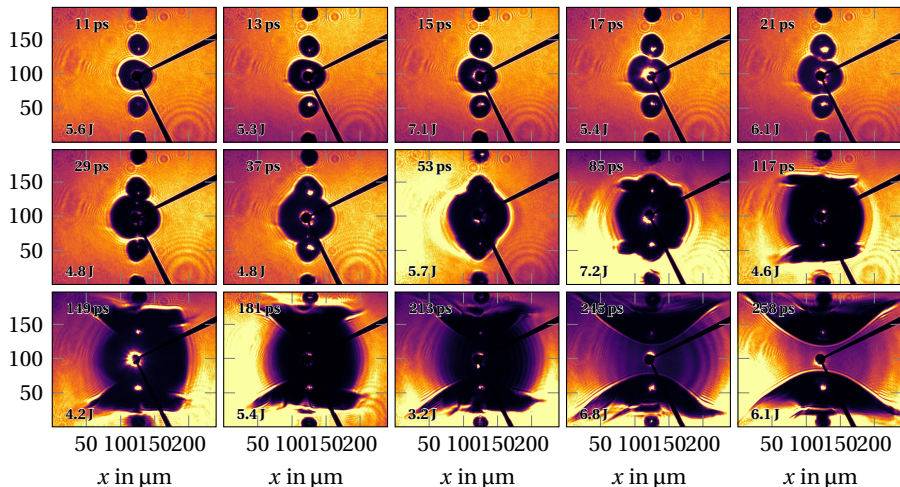


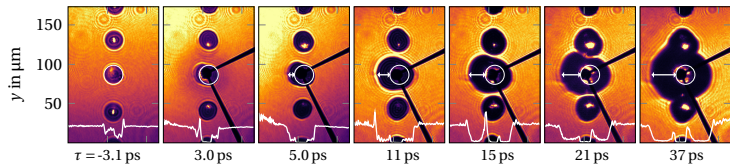


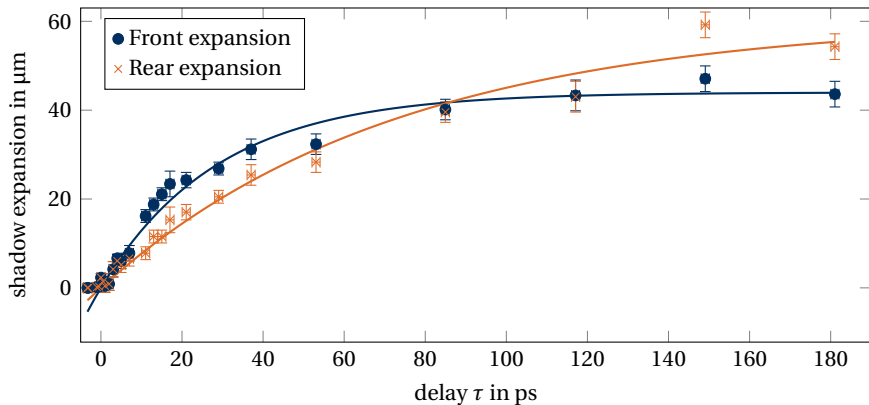
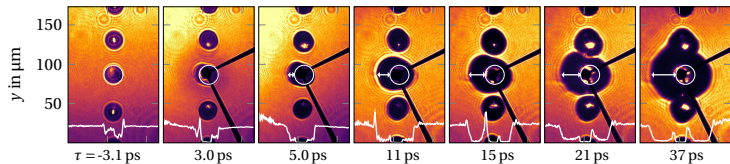
Shadowgraphy of the plasma evolution



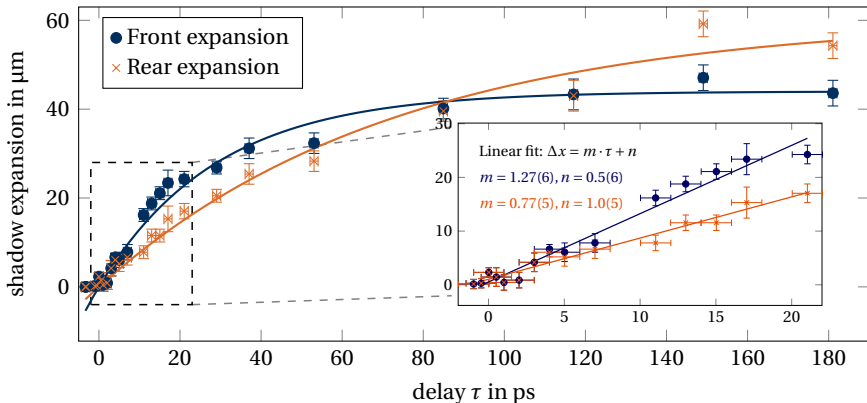
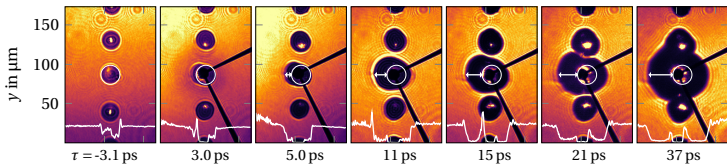
Shadowgraphy of the plasma evolution



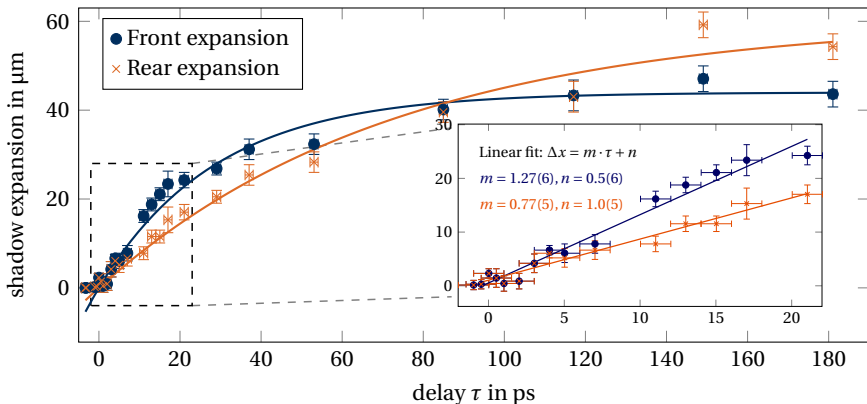




Estimation of plasma expansion velocity



$\tau = 3-20 \text{ ps} \quad v_{\text{front}} = (1.27 \pm 0.06) \mu\text{m}/\text{ps}, \quad v_{\text{rear}} = (0.77 \pm 0.05) \mu\text{m}/\text{ps}$
 $\tau = 80-190 \text{ ps} \quad v_{\text{front}} = (0.09 \pm 0.03) \mu\text{m}/\text{ps}, \quad v_{\text{rear}} = (0.22 \pm 0.05) \mu\text{m}/\text{ps}.$



$$\begin{aligned} \tau = 3-20 \text{ ps} & \quad v_{\text{front}} = (1.27 \pm 0.06) \mu\text{m/ps}, & v_{\text{rear}} = (0.77 \pm 0.05) \mu\text{m/ps} \\ \tau = 80-190 \text{ ps} & \quad v_{\text{front}} = (0.09 \pm 0.03) \mu\text{m/ps}, & v_{\text{rear}} = (0.22 \pm 0.05) \mu\text{m/ps}. \end{aligned}$$

Table: Comparison of shadowgraphy plasma expansion experiments with targets for different peak intensities.

experiment	target	intensity	λ	v_{front}
Becker ¹	H ₂ O droplets	10^{16} W/cm^2	800 nm	$0.38 \mu\text{m/ps}$
This thesis	H ₂ O droplets	$4 \cdot 10^{19} \text{ W/cm}^2$	1030 nm	$1.3 \mu\text{m/ps}$
Bernert et al. ²	hydrogen jet	$5 \cdot 10^{21} \text{ W/cm}^2$	800 nm	$23 \mu\text{m/ps}$

¹ “Characterization of laser-driven proton acceleration with contrast-enhanced laser pulses”

² “Off-harmonic optical probing of high intensity laser plasma expansion dynamics in solid density hydrogen jets”

1. Fundamentals

2. Probing relativistic laser-plasma interactions

3. SPM-based laser pulse generation

- Self phase modulation (SPM)
- Single-pass SPM-based laser pulses
- Multi-pass SPM-based laser pulses
- Impact of pulse steepening

Optical Kerr effect (intensity dependent refractive index)

$$n(t) = n_0 + n_2 I(t), \quad \text{where} \quad n_2 = \frac{3\chi^{(3)}}{4\epsilon_0 n_0^2 c}.$$

Optical Kerr effect (intensity dependent refractive index)

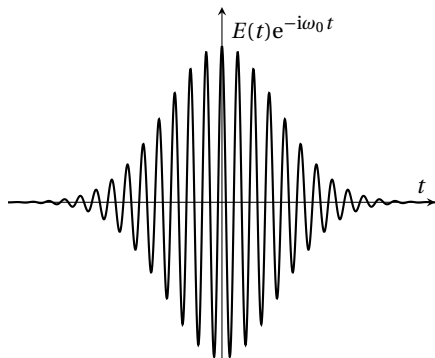
$$n(t) = n_0 + n_2 I(t), \quad \text{where} \quad n_2 = \frac{3\chi^{(3)}}{4\epsilon_0 n_0^2 c}.$$

Change in the phase: $\Phi_{\text{NL}}(t) = n_2 I(t) k_0 L.$

Optical Kerr effect (intensity dependent refractive index)

$$n(t) = n_0 + n_2 I(t), \quad \text{where} \quad n_2 = \frac{3\chi^{(3)}}{4\epsilon_0 n_0^2 c}.$$

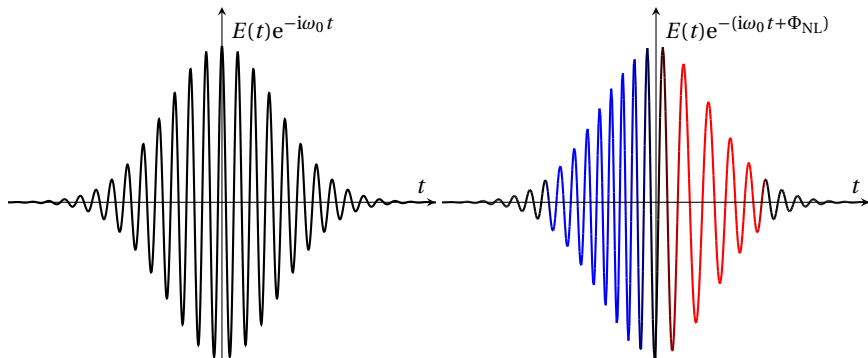
Change in the phase: $\Phi_{\text{NL}}(t) = n_2 I(t) k_0 L$.



Optical Kerr effect (intensity dependent refractive index)

$$n(t) = n_0 + n_2 I(t), \quad \text{where} \quad n_2 = \frac{3\chi^{(3)}}{4\varepsilon_0 n_0^2 c}.$$

Change in the phase: $\Phi_{\text{NL}}(t) = n_2 I(t) k_0 L$.



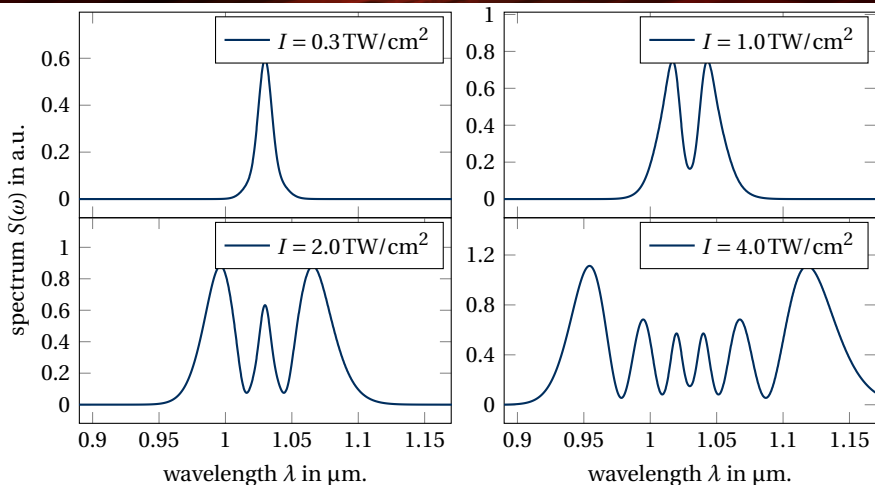
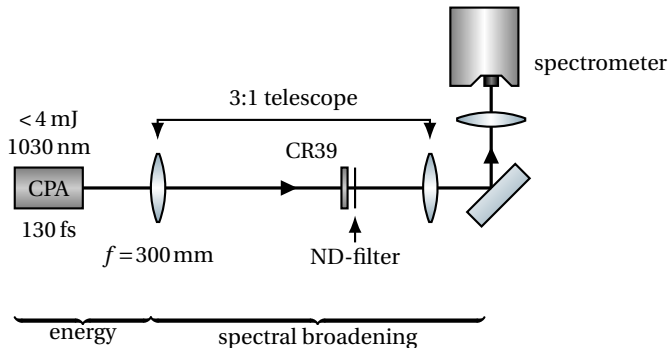
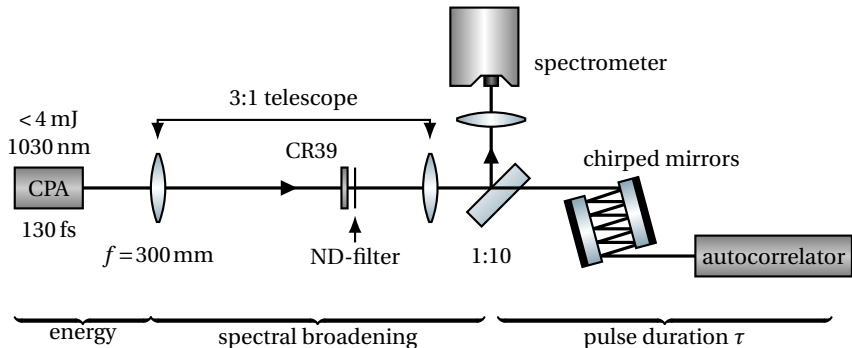


Fig.: Simulated spectra of a laser pulse with $\tau = 130$ fs (FWHM) propagating through a 1 mm thick foil of CR-39 at different intensities.

- **Pulse duration** $\tau < 130$ fs
⇒ high temporal resolution
- **Wavelength** $\lambda = 800 - 1000$ nm
⇒ No overlap with fundamental (1030 nm) and SH (515 nm) of the POLARIS laser
- **Spectral bandwidth** $\Delta\lambda > 100$ nm
⇒ Support femtosecond pulse durations
- **Energy** $E > 10 \mu\text{J}$
⇒ proper illumination, no target ionization





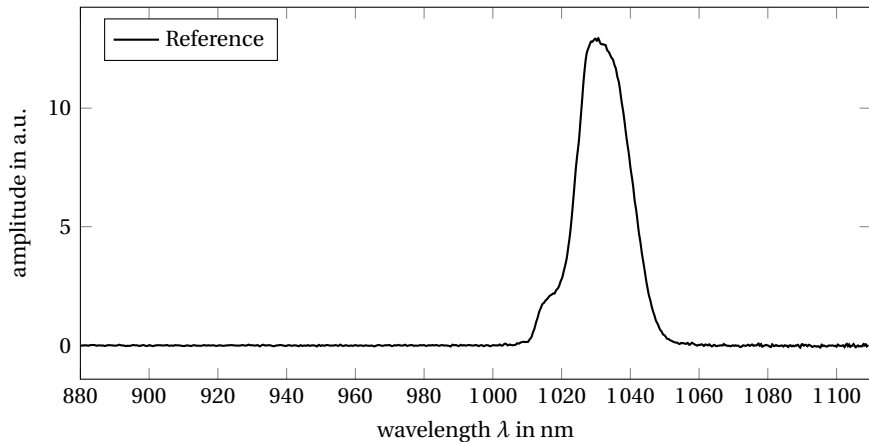


Fig.: Measurement of SPM-induced spectra in CR-39 for a pulse energy of 0.5 mJ at different intensities by varying the sample distance to the focus position. A reference spectrum in air was taken at a lower pulse energy of 0.1 mJ.

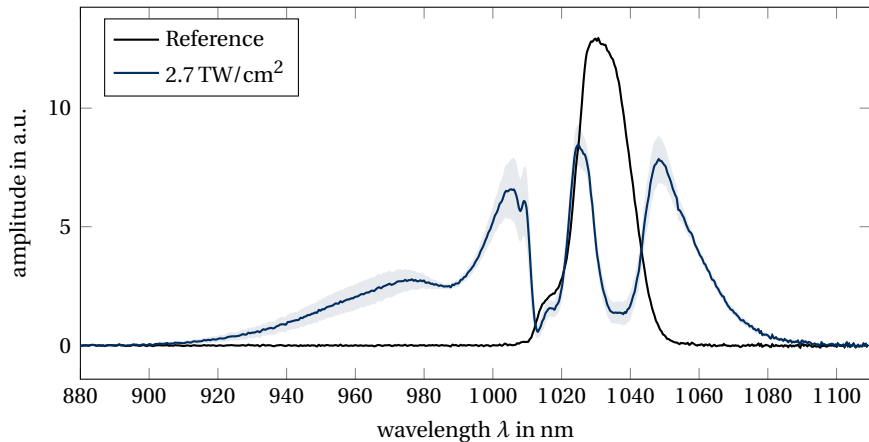


Fig.: Measurement of SPM-induced spectra in CR-39 for a pulse energy of 0.5 mJ at different intensities by varying the sample distance to the focus position. A reference spectrum in air was taken at a lower pulse energy of 0.1 mJ.

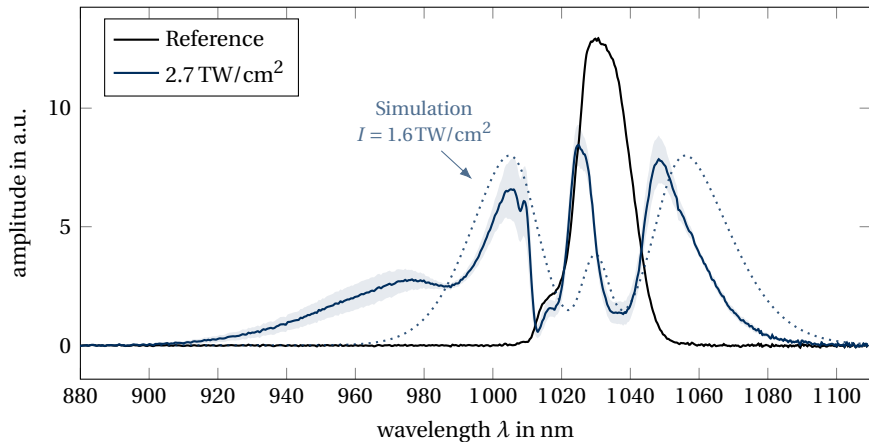


Fig.: Measurement of SPM-induced spectra in CR-39 for a pulse energy of 0.5 mJ at different intensities by varying the sample distance to the focus position. A reference spectrum in air was taken at a lower pulse energy of 0.1 mJ.

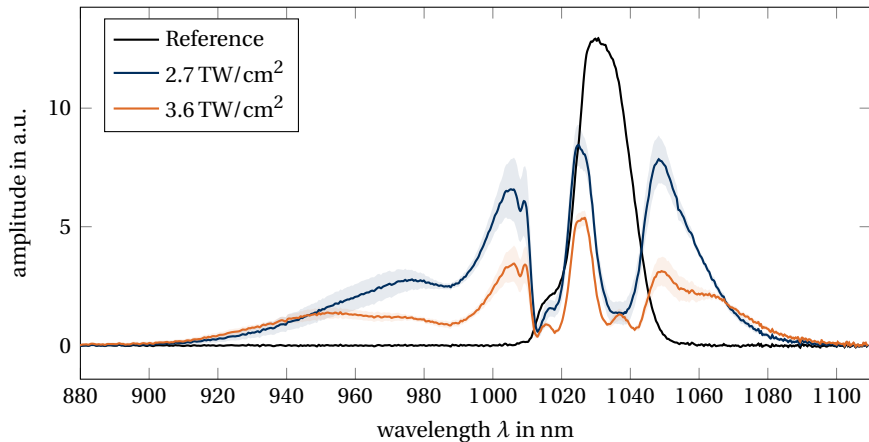


Fig.: Measurement of SPM-induced spectra in CR-39 for a pulse energy of 0.5 mJ at different intensities by varying the sample distance to the focus position. A reference spectrum in air was taken at a lower pulse energy of 0.1 mJ.

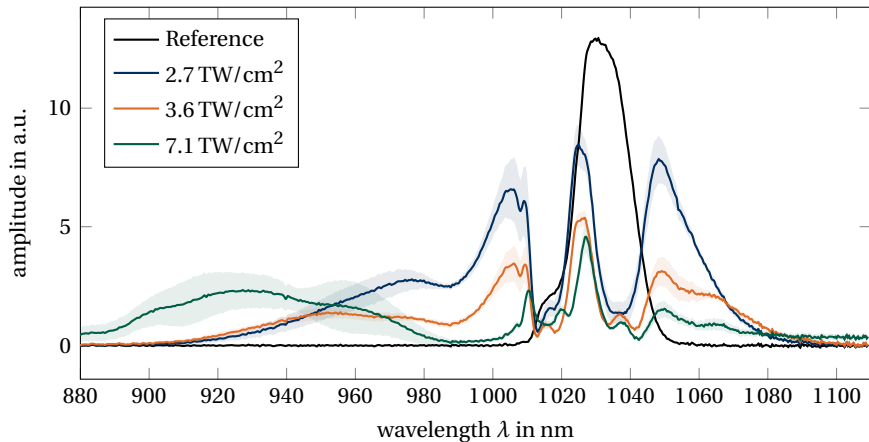


Fig.: Measurement of SPM-induced spectra in CR-39 for a pulse energy of 0.5 mJ at different intensities by varying the sample distance to the focus position. A reference spectrum in air was taken at a lower pulse energy of 0.1 mJ.

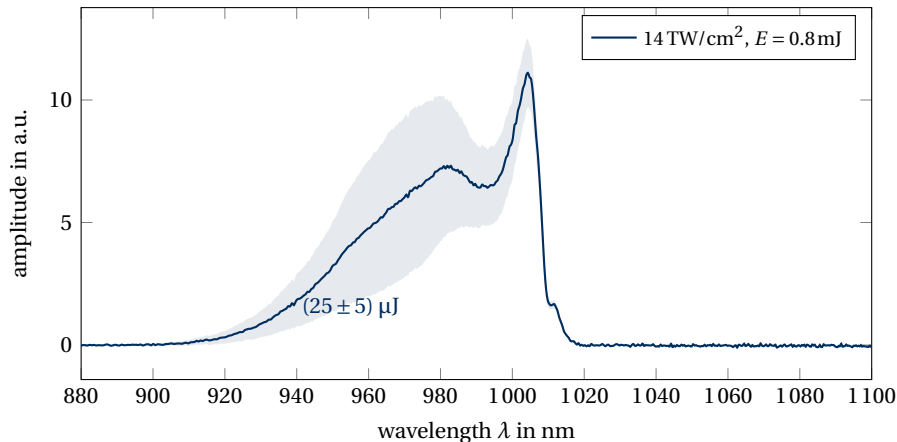


Fig.: Measurement of SPM in CR-39 performed at low pressure (5 mbar) with a short-pass filter cutting at 1010 nm.

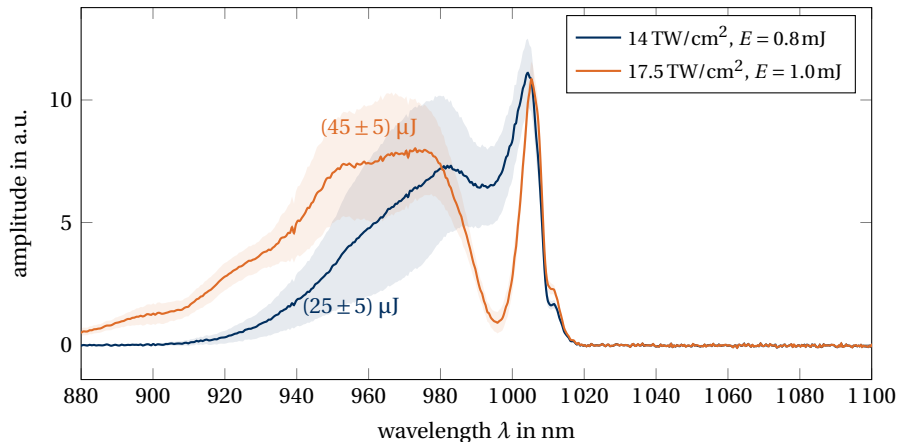


Fig.: Measurement of SPM in CR-39 performed at low pressure (5 mbar) with a short-pass filter cutting at 1010 nm.

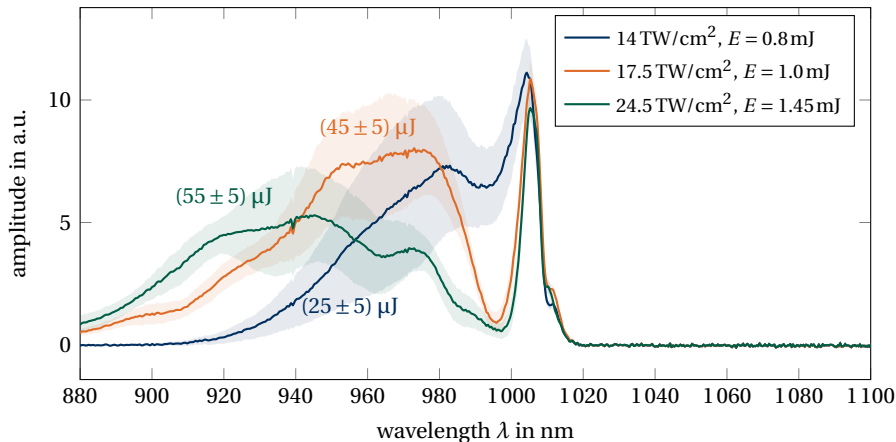
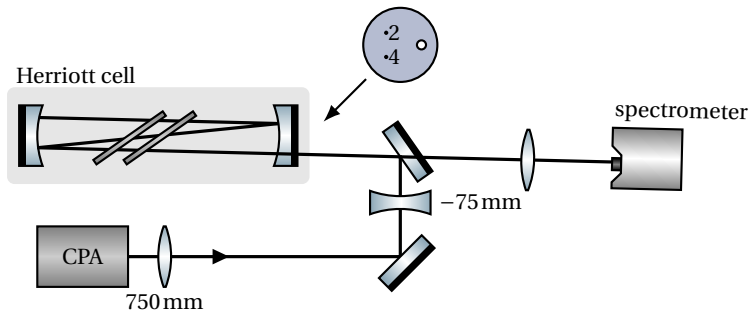


Fig.: Measurement of SPM in CR-39 performed at low pressure (5 mbar) with a short-pass filter cutting at 1010 nm.



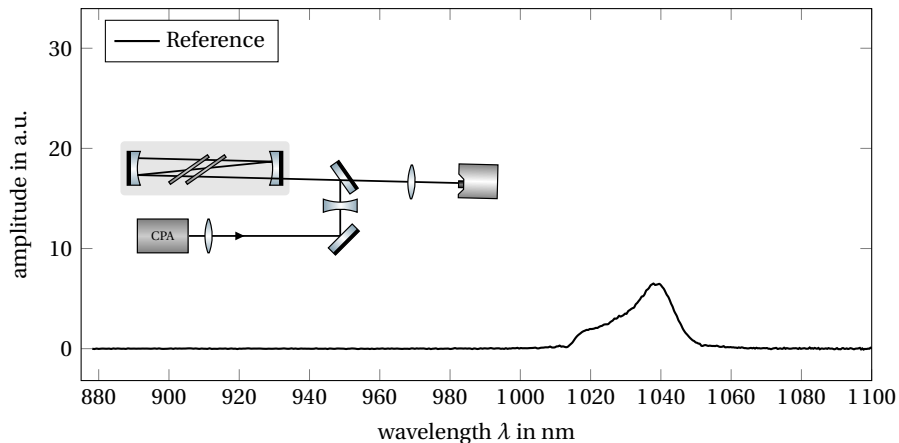


Fig.: Measured Spectra in two samples of 1.1 mm CR-39 placed in the Brewster angle into a MPC of length $L = 10$ cm.

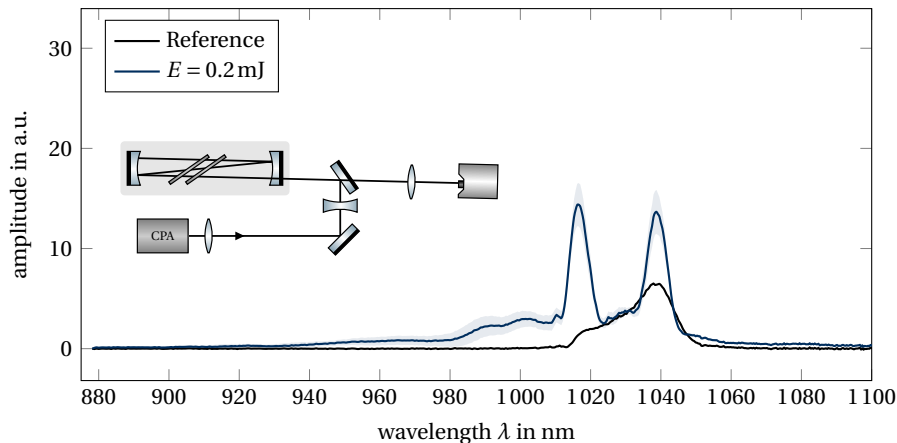


Fig.: Measured Spectra in two samples of 1.1 mm CR-39 placed in the Brewster angle into a MPC of length $L = 10$ cm.

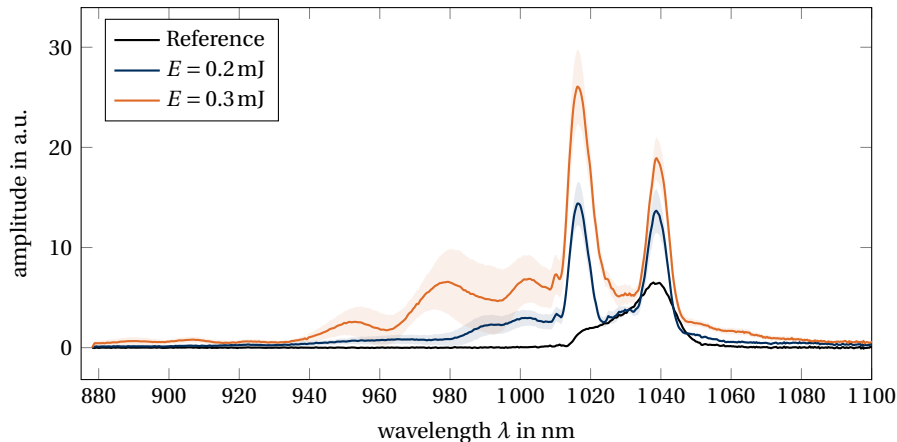


Fig.: Measured Spectra in two samples of 1.1 mm CR-39 placed in the Brewster angle into a MPC of length $L = 10$ cm.

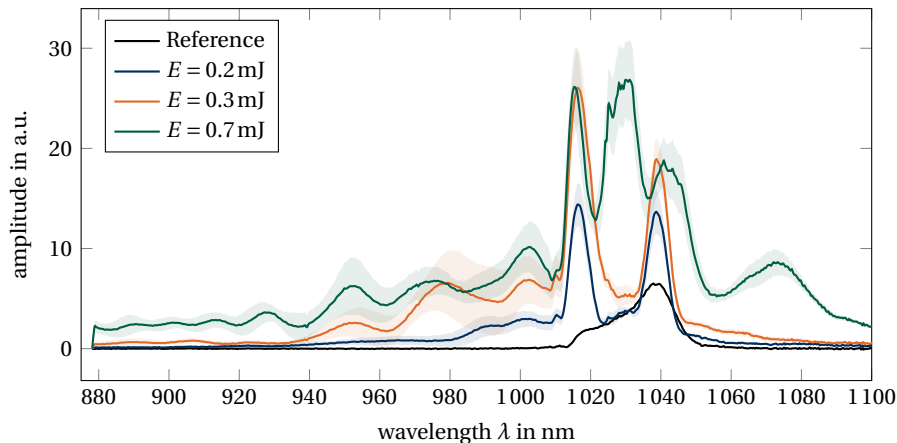


Fig.: Measured Spectra in two samples of 1.1 mm CR-39 placed in the Brewster angle into a MPC of length $L = 10$ cm.

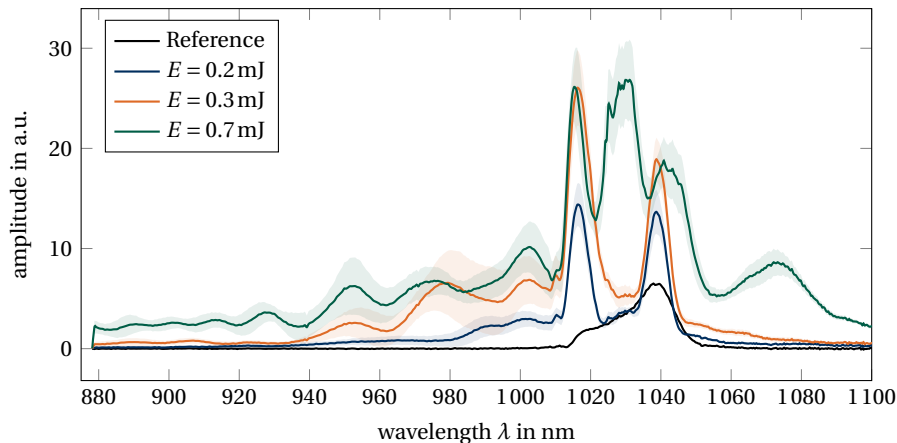


Fig.: Measured Spectra in two samples of 1.1 mm CR-39 placed in the Brewster angle into a MPC of length $L = 10$ cm.

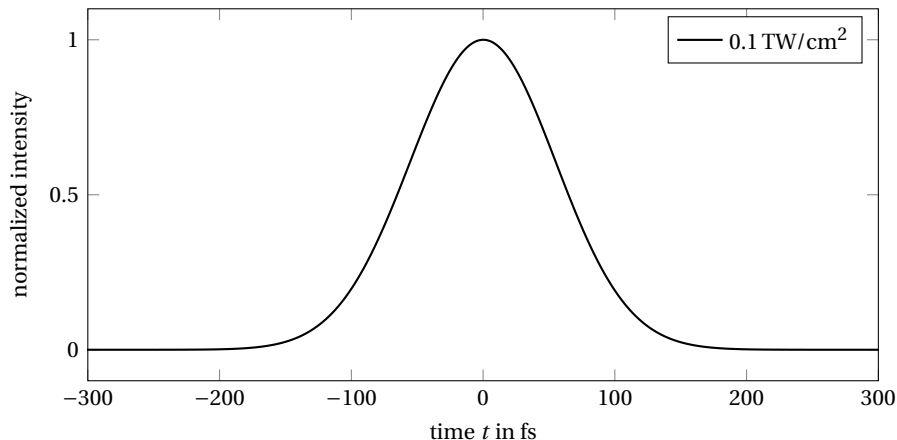


Fig.: Simulated temporal profiles of the pulse intensities for a 1.1 mm thick sheet of CR-39 with an initial pulse length of 130 fs (FWHM) for different intensities.

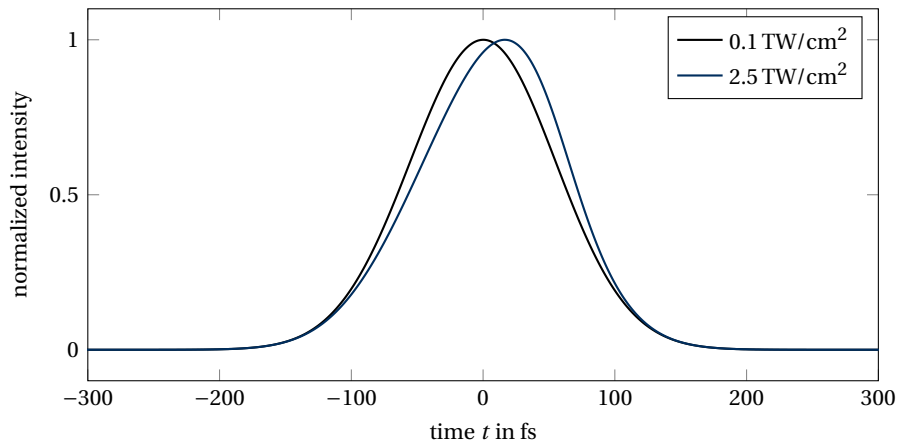


Fig.: Simulated temporal profiles of the pulse intensities for a 1.1 mm thick sheet of CR-39 with an initial pulse length of 130 fs (FWHM) for different intensities.

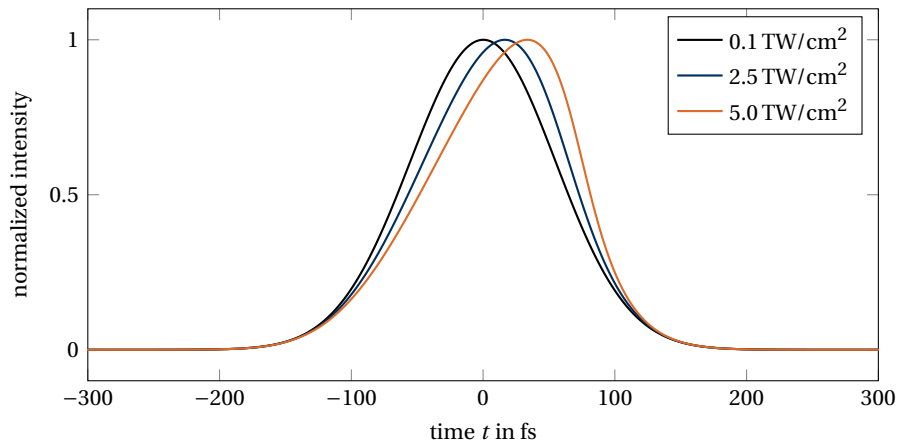


Fig.: Simulated temporal profiles of the pulse intensities for a 1.1 mm thick sheet of CR-39 with an initial pulse length of 130 fs (FWHM) for different intensities.

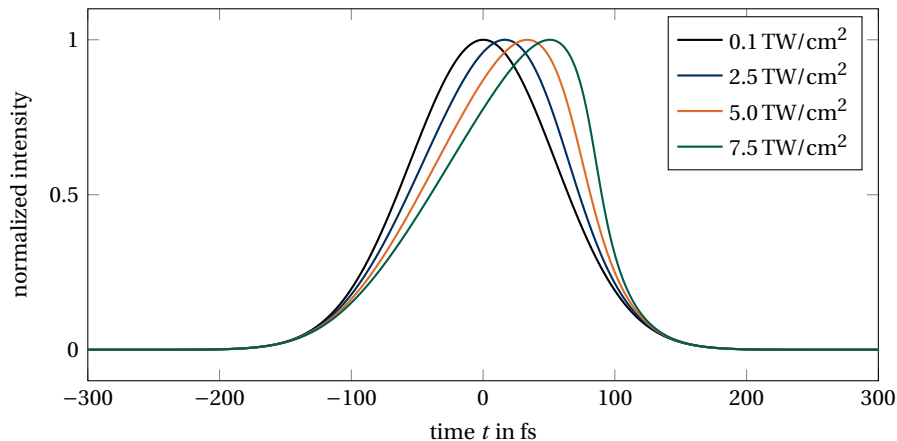


Fig.: Simulated temporal profiles of the pulse intensities for a 1.1 mm thick sheet of CR-39 with an initial pulse length of 130 fs (FWHM) for different intensities.

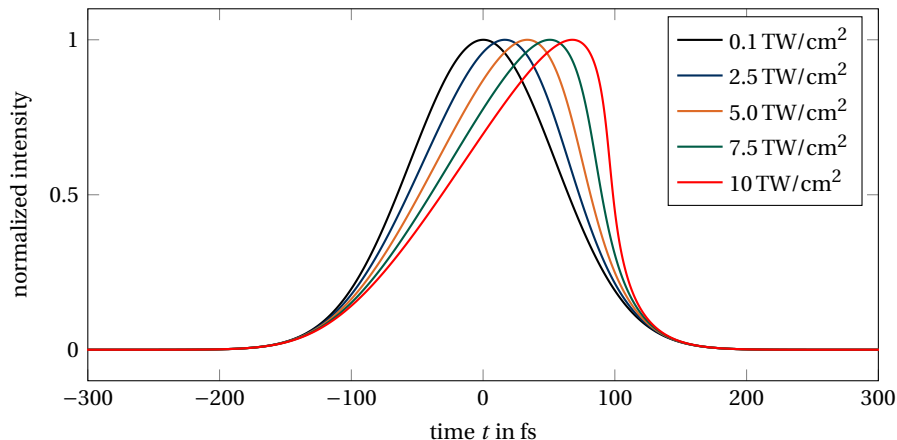


Fig.: Simulated temporal profiles of the pulse intensities for a 1.1 mm thick sheet of CR-39 with an initial pulse length of 130 fs (FWHM) for different intensities.

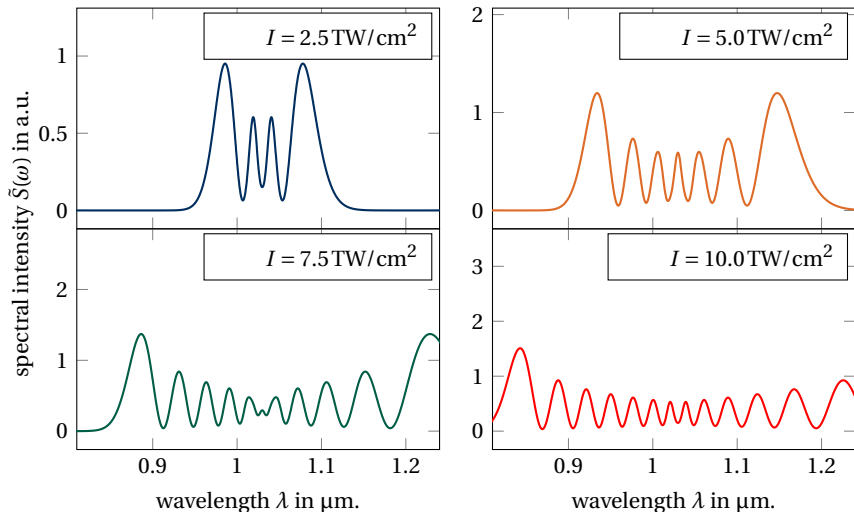


Fig.: Calculated spectra of an ultrashort pulse propagating through CR-39.

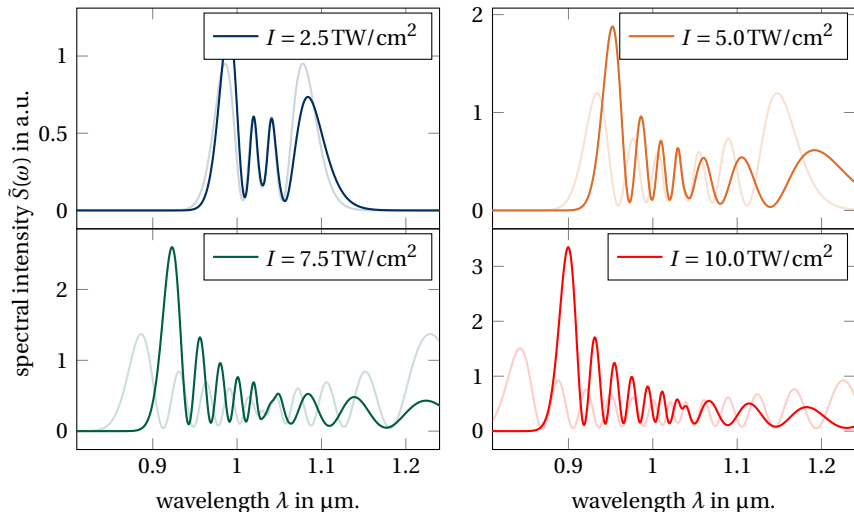


Fig.: Calculated spectra of an ultrashort pulse propagating through CR-39.

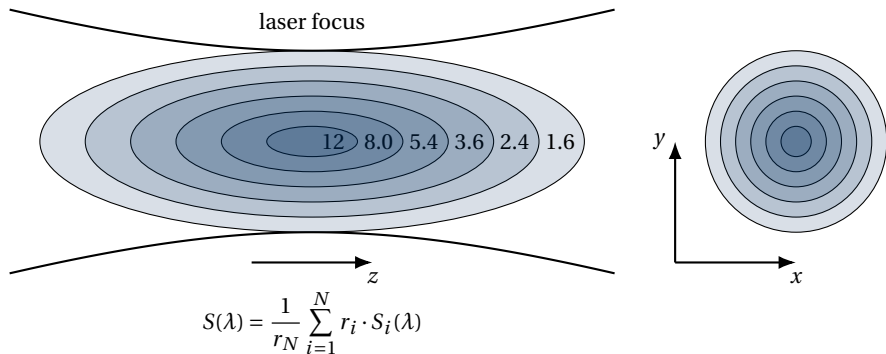


Fig.: In a simplified model, the intensity distribution can be approximated as rings of constant thickness. The corresponding intensity drops according to the Gaussian distribution exponentially.

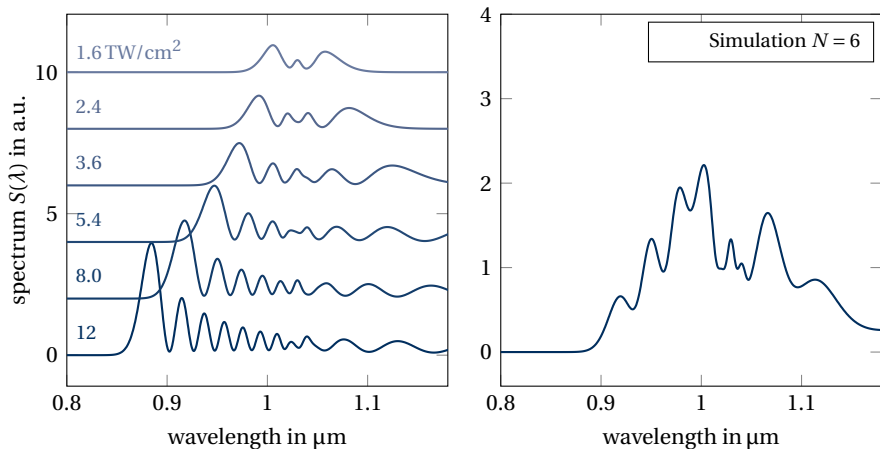


Fig.: Left: Calculated spectra including the effect of pulse steepening in CR-39.
Right: Focal volume averaging for a transverse Gaussian intensity distribution.

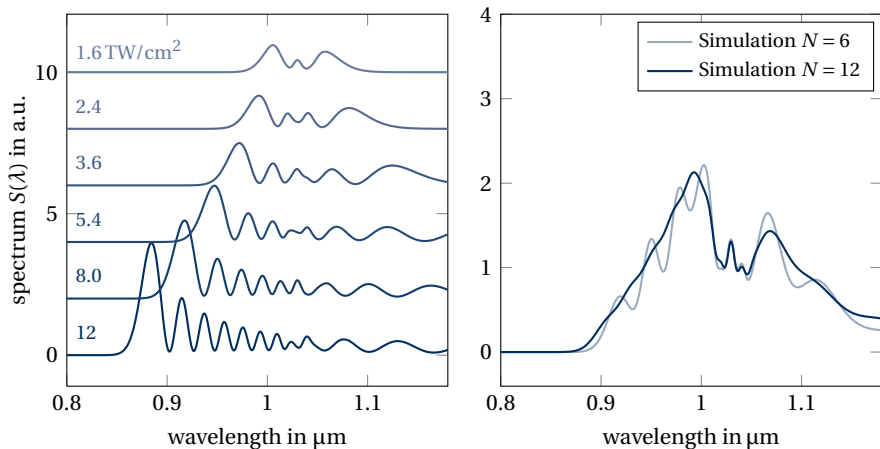


Fig.: Left: Calculated spectra including the effect of pulse steepening in CR-39.
Right: Focal volume averaging for a transverse Gaussian intensity distribution.

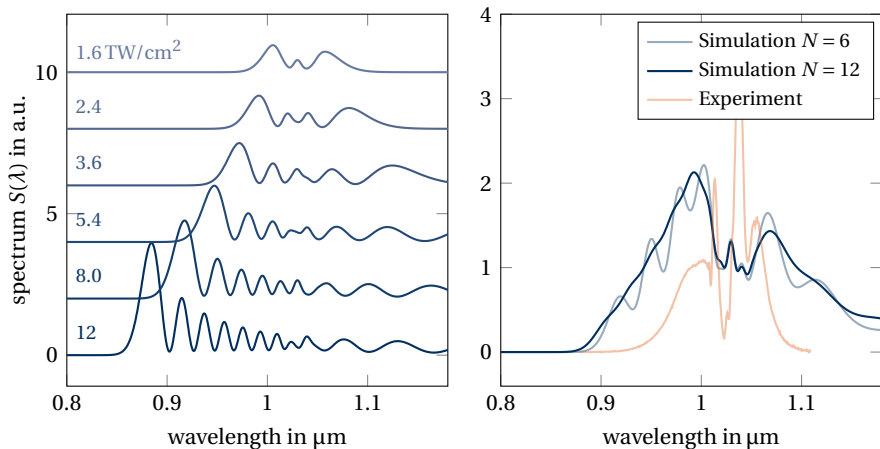


Fig.: Left: Calculated spectra including the effect of pulse steepening in CR-39.
Right: Focal volume averaging for a transverse Gaussian intensity distribution.

Probing of laser-microdroplet interaction

- characterization of the NOPA setup
- development: imaging diagnostics + coronagraph
- plasma evolution + plasma expansion velocity

Development of a new probe pulse laser source

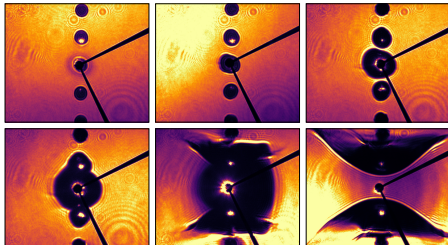
- Single-pass SPM \Rightarrow 50 μ J pulses
- Multi-pass SPM \Rightarrow lower intensities
- Simulation: Pulse steepening + focal averaging

A special thanks to my supervisors:

- Prof. Dr. Malte Kaluza
- Dr. Yasmina Azamoum

I would also like to thank:

- Till Weickhardt
- Mathis Nolte
- Marco Hellwing
- Dr. Georg Becker
- Dr. Marco Hornung
- Dr. Matthew Schwab



Thank you for your attention.

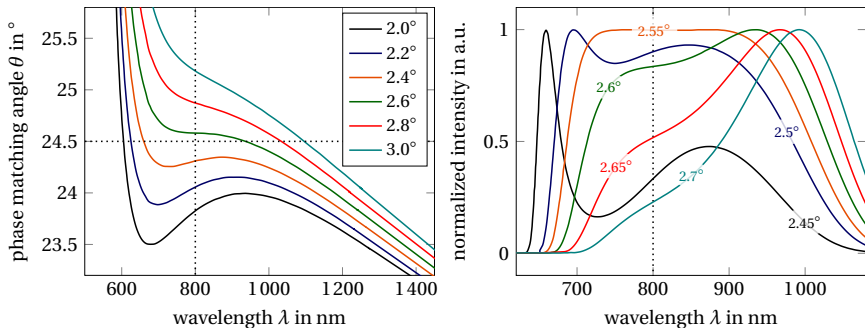


Fig.: Left: Type-1 phase-matching angle θ as a function of the signal wavelength λ_s for different pump-signal angles α with a pump wavelength at 515 nm. Right: Calculated gain spectra for a phase matching angle of $\theta = 24.5^\circ$ as a function of signal wavelength λ_s for different pump-signal angles α .

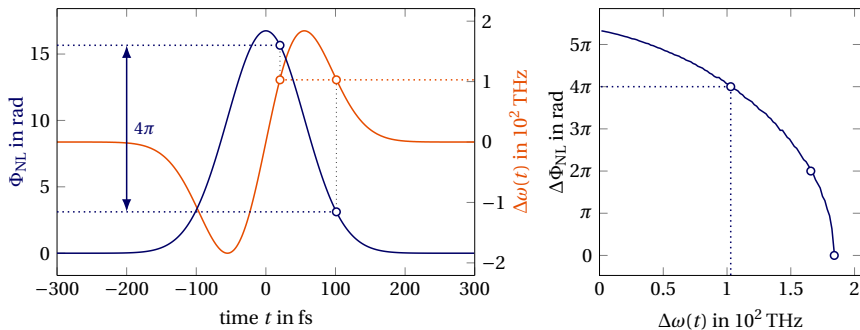


Fig.: Left: Nonlinear phase (blue) for a pulse with $\tau = 130$ fs (FWHM) of peak intensity $I_0 = 4$ TW/cm² for 1 mm of CR-39. The instantaneous frequency change $\Delta\omega = -\frac{\partial\Phi_{NL}}{\partial t}$ is shown in orange. Right: Calculated phase difference as a function of instantaneous frequency change.

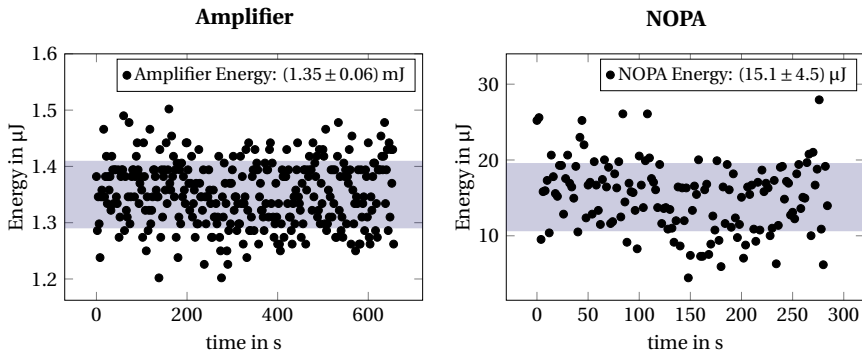


Fig.: Stability measurement of the amplifier and NOPA energy over a duration of several minutes.

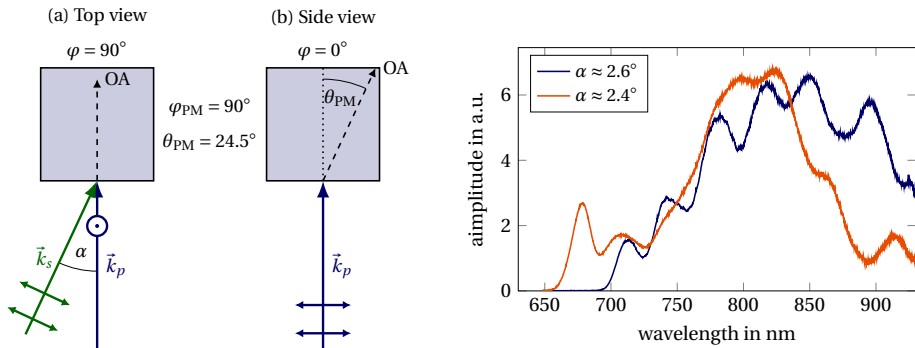
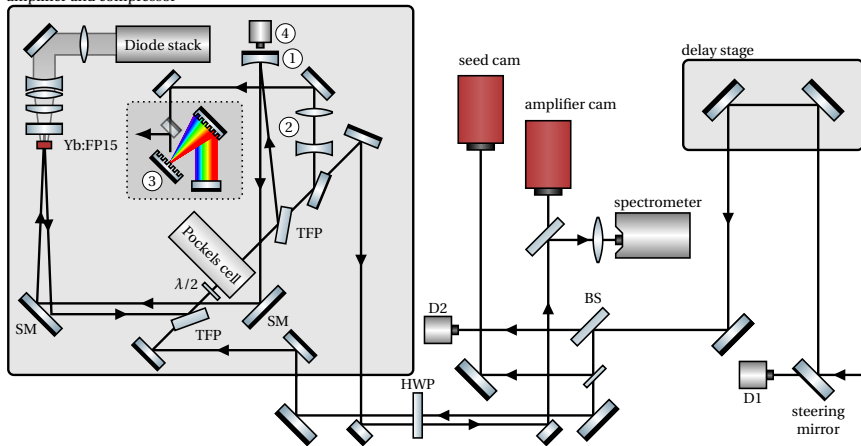


Fig.: Left: Geometry of the WLC (signal) and SH (pump) beams in the NOPA crystal (Type I o-o-e phase matching process). Right: NOPA spectra for different non-collinear angles α . The orange curve indicates that the spectral bandwidth can be reduced by an adjustment of the non-collinear angle and subsequent optimization of the temporal delay τ .

amplifier and compressor



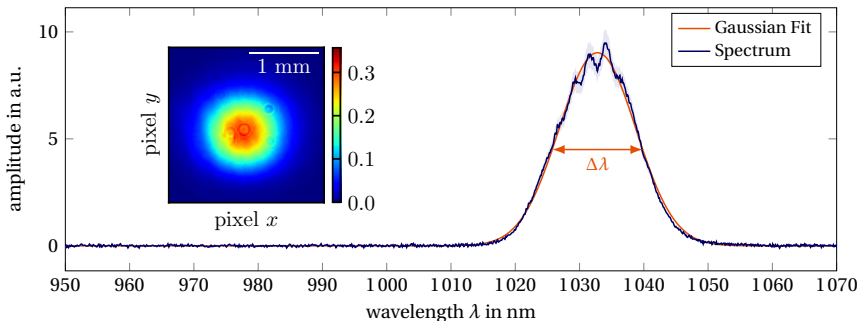


Fig.: Measured spectrum of the Amplifier. The graph shows the averaged spectrum for 80 consecutive shots. The standard deviation is indicated with the shaded area. The central wavelength $\lambda_0 = 1032.8$ nm, and width $\Delta\lambda = 14.1$ nm (FWHM) were determined using a Gaussian fit (orange). Inset: spatial profile of the amplified pulse. The beam radius was determined to be $w = (0.77 \pm 0.02)$ mm ($1/e^2$).

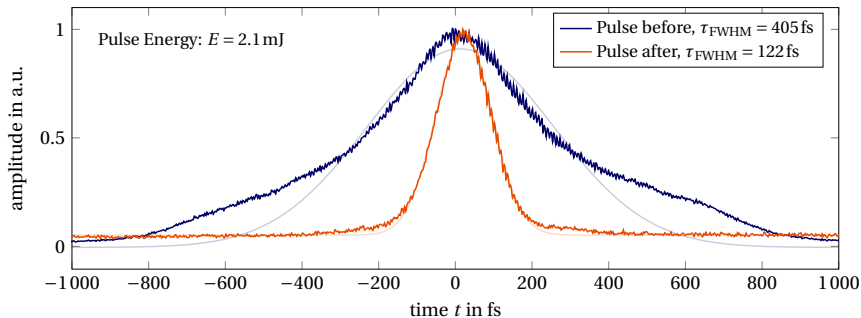


Fig.: Measured pulse length at the compressor output before (blue) and after (orange) the realignment of the stretcher. The autocorrelation function was measured by TOPAG ASF-15 single shot autocorrelator. The desaturated curves show a Gaussian fit to the measured data used to determine the pulse length.

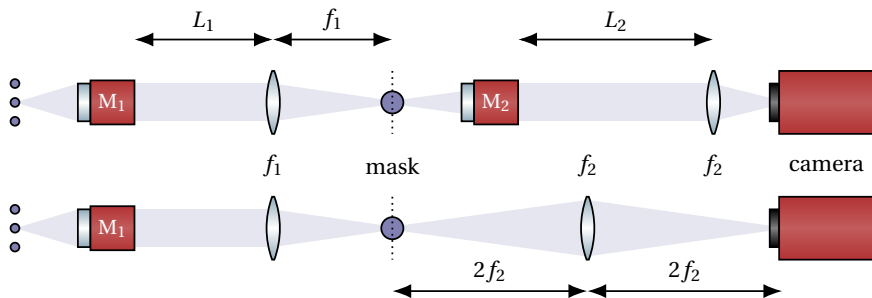


Fig.: Two different designs for a coronagraph imaging setup. Top: The droplets are magnified and imaged to the position of the coronagraph mask using an aberration corrected microscope objective M_1 and a tube lens (f_1) placed in a distance L_1 to the objective. The droplet with the blocked center is then re-imaged into a camera with another combination of objective and lens f_2 placed in a distance L_2 . Bottom: The second imaging of the droplet into the camera is done by a single lens with a magnification of 1.

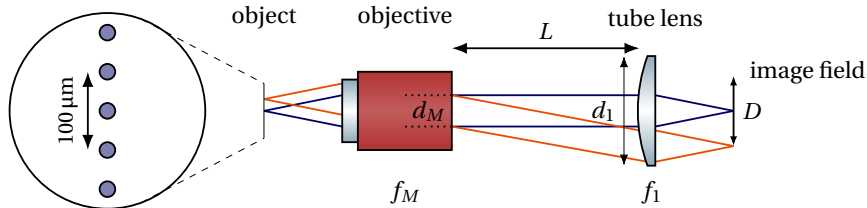


Fig.: Supplemental drawing for the placement of objective and tube lens. The maximum distance L for a desired image field D can be derived geometrically from the tube lens diameter d_1 and objective aperture d_M . The required field of view with a size of $300\ \mu\text{m}$ is shown on the left.

Self-focusing distance z_{sf} :

$$z_{\text{sf}} = \frac{2n_0 w_0^2}{\lambda_0} \frac{1}{\sqrt{P/P_{\text{cr}} - 1}} \quad \text{with} \quad P_{\text{crit}} \approx \frac{\lambda_0^2}{8n_0 n_2} = 1.42 \text{ MW} \quad \text{at } \lambda_0 = 1030 \text{ nm},$$

Table: Self focusing distance z_{sf} and pulse width w_0 for different pulse energies ($\tau = 130$ fs) calculated at a constant peak intensity $I = 10 \text{ TW/cm}^2$.

Energy [mJ]	0.1	0.5	1	2	5
pulse radius w_0 [μm]	48	107	152	214	339
self-focusing distance z_{sf} [mm]	0.30	0.66	0.94	1.32	2.09

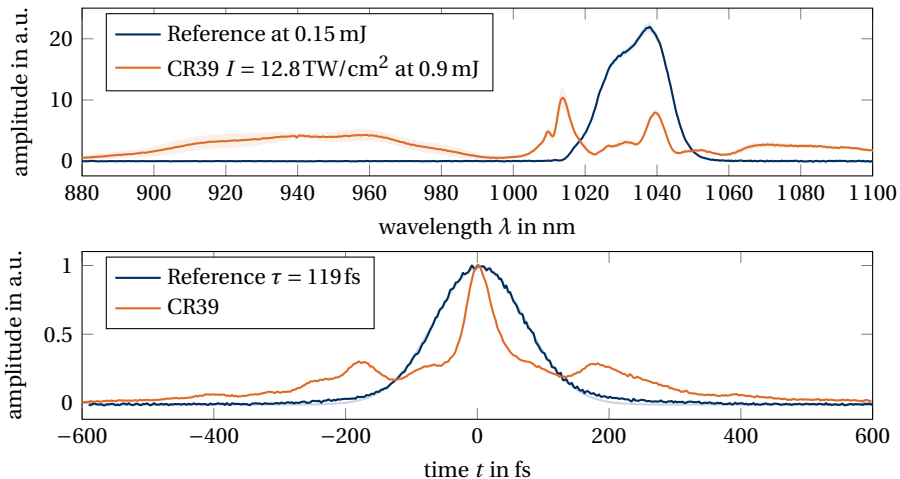
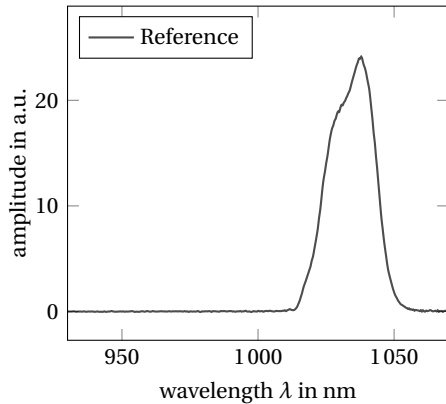
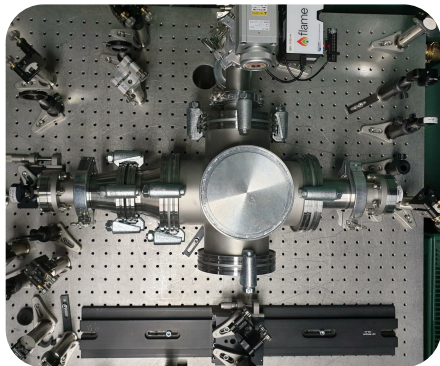
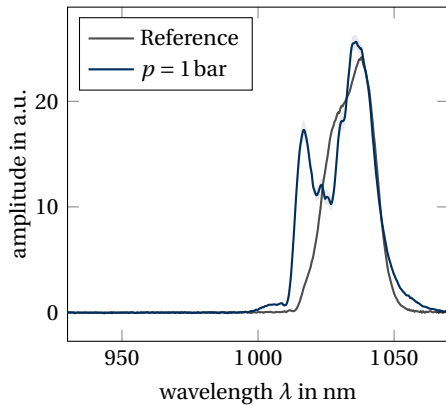
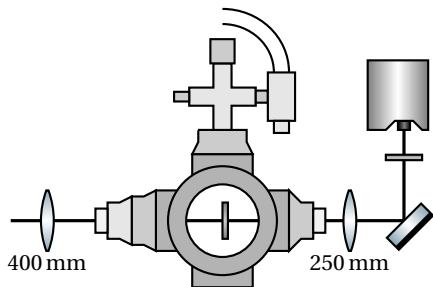
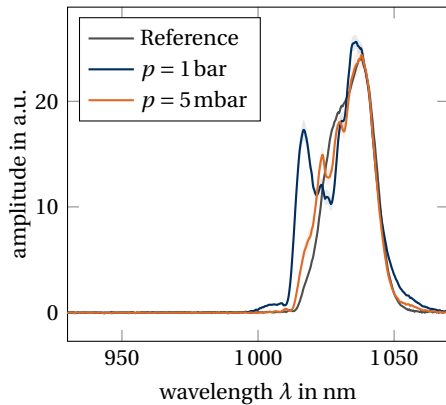
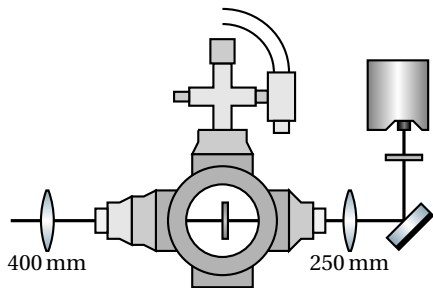


Fig.: Top: Spectra of a reference pulse (blue) and for pulses experiencing SPM in CR-39 (orange) recorded simultaneously with an autocorrelation measurement. Bottom: Measured Autocorrelation functions.







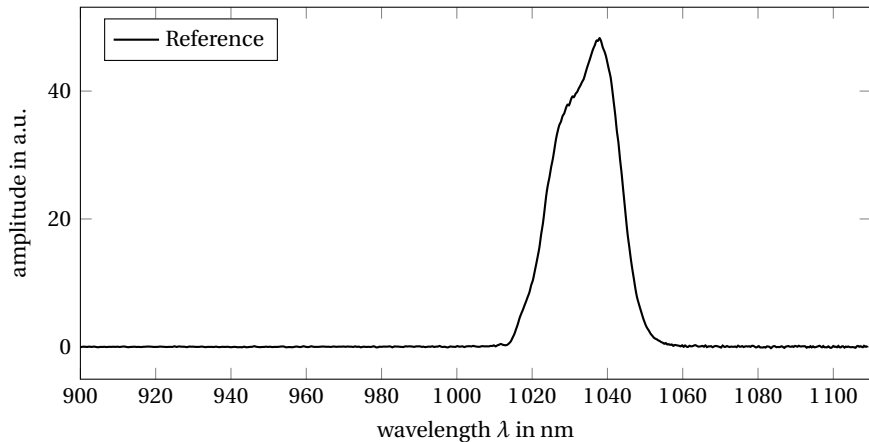


Fig.: Measurement of SPM in CR-39 performed at low pressure (5 mbar). Different pulse energies were used at a constant distance (2 cm) to the laser focus (focusing lens $f = 400$ mm).

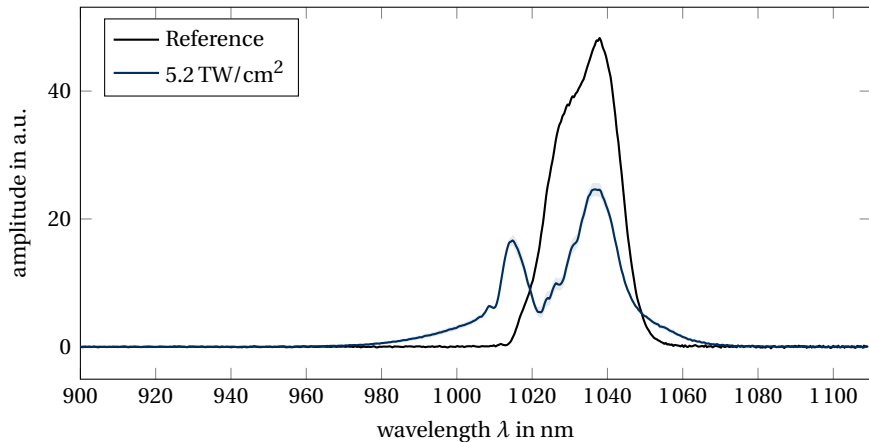


Fig.: Measurement of SPM in CR-39 performed at low pressure (5 mbar). Different pulse energies were used at a constant distance (2 cm) to the laser focus (focusing lens $f = 400$ mm).

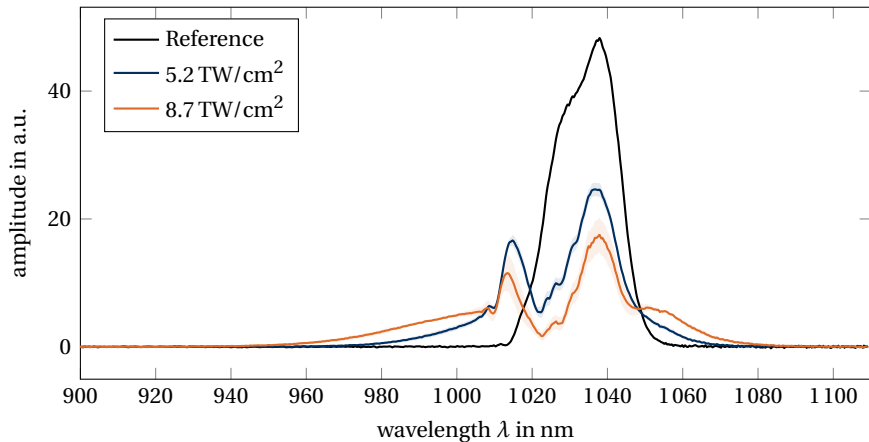


Fig.: Measurement of SPM in CR-39 performed at low pressure (5 mbar). Different pulse energies were used at a constant distance (2 cm) to the laser focus (focusing lens $f = 400 \text{ mm}$).

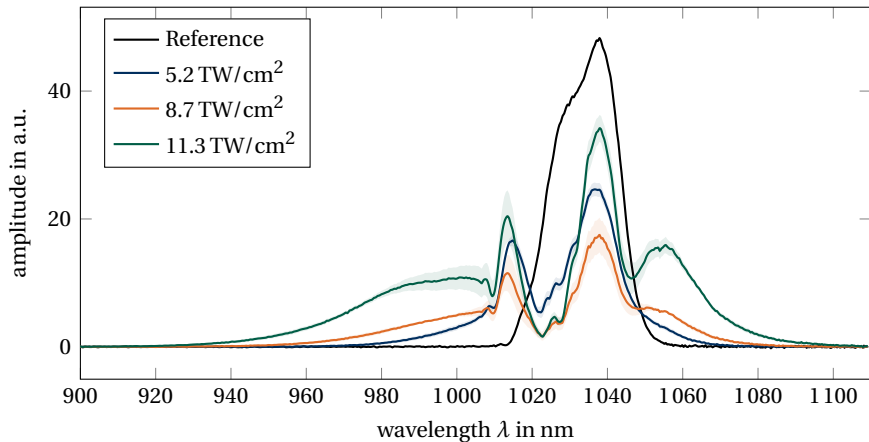


Fig.: Measurement of SPM in CR-39 performed at low pressure (5 mbar). Different pulse energies were used at a constant distance (2 cm) to the laser focus (focusing lens $f = 400$ mm).

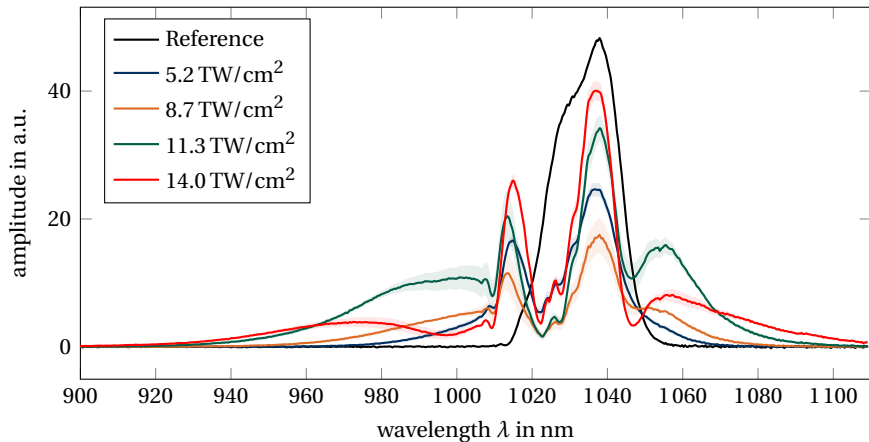
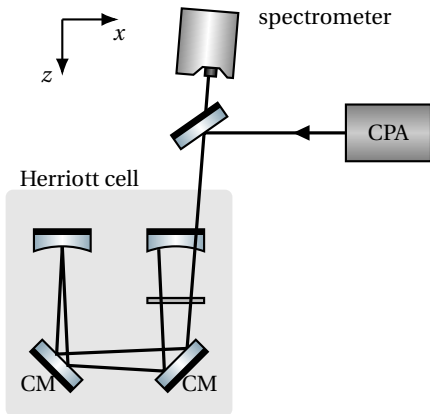
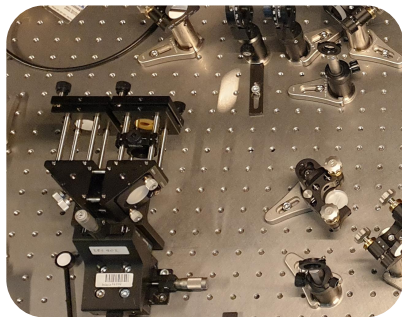
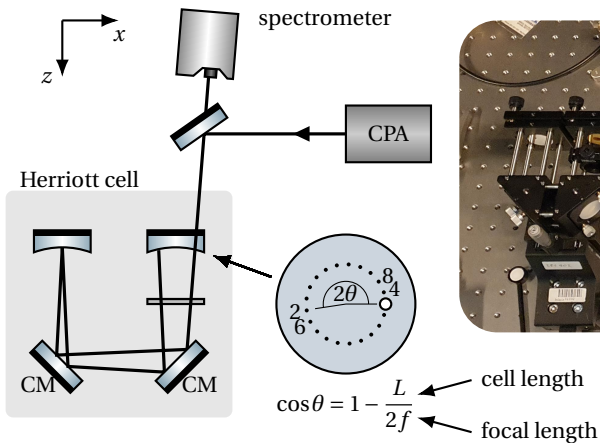
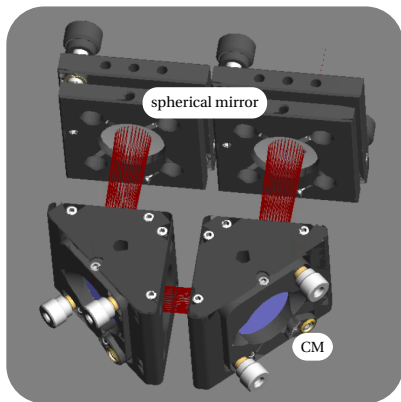
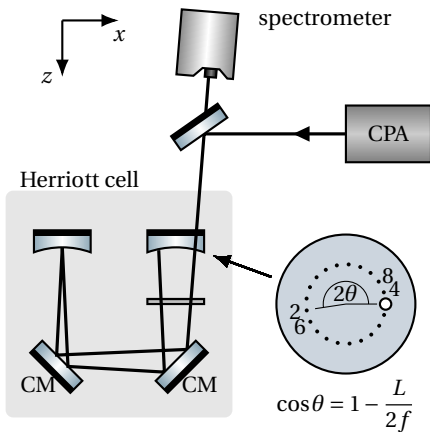


Fig.: Measurement of SPM in CR-39 performed at low pressure (5 mbar). Different pulse energies were used at a constant distance (2 cm) to the laser focus (focusing lens $f = 400$ mm).







Wave equation of a plane wave in a nonlinear, inversion symmetric medium with slowly varying envelope approximation

$$\left(\frac{\partial^2}{\partial z^2} - \frac{n_0^2}{c^2} \frac{\partial^2}{\partial t^2} \right) \mathbf{E} = \frac{1}{c^2 \epsilon_0} \frac{\partial^2}{\partial t^2} \mathbf{P}^{(3)} \quad \text{with} \quad \mathbf{P}^{(3)} = \frac{3}{4} \epsilon_0 \chi^{(3)} |\mathbf{E}|^2 \mathbf{E}.$$

Without slowly varying envelope approximation: $(\mathbf{E}(t) = \mathcal{E} e^{i(kz - \omega t)})$

$$\left(\frac{\partial}{\partial z} + \frac{n_0}{c} \frac{\partial}{\partial t} \right) \mathcal{E} + \frac{1}{2ik} \left(\frac{\partial^2}{\partial z^2} - \frac{n_0^2}{c^2} \frac{\partial^2}{\partial t^2} \right) \mathcal{E} = \frac{1}{2ik} \frac{3\omega_0^2}{4c^2} \left(\frac{1}{\omega_0^2} \frac{\partial^2}{\partial t^2} - \frac{2i}{\omega_0} \frac{\partial}{\partial t} - 1 \right) \chi^{(3)} |\mathcal{E}|^2 \mathcal{E}.$$

Further simplification and splitting of amplitude $|\mathcal{E}|$ and phase Φ yields

$$\left[\frac{\partial}{\partial z} + \frac{n_0}{c} \left(1 + \frac{3\tilde{n}_2}{n_0} |\mathcal{E}|^2 \right) \frac{\partial}{\partial t} \right] |\mathcal{E}| = 0,$$

$$\left[\frac{\partial}{\partial z} + \frac{n_0}{c} \left(1 + \frac{\tilde{n}_2}{n_0} |\mathcal{E}|^2 \right) \frac{\partial}{\partial t} \right] \Phi = \frac{\tilde{n}_2 \omega_0}{c} |\mathcal{E}|^2.$$